

Lecture 2

①

1) Recap. from lecture 1

$$a) U(\Lambda) |k, \lambda\rangle = e^{i\Delta\theta(\mathcal{L}^{-1}(\Lambda k) \wedge \mathcal{L}(k))} |\Lambda k, \lambda\rangle$$

element of
the little group

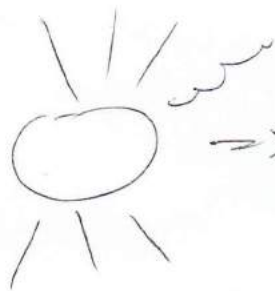
$\mathcal{L}(k)$ - standard transformation

$$\vec{k}^\wedge \xrightarrow{\mathcal{L}} k^\wedge = (|\vec{k}|, \vec{k})$$

\uparrow
($\kappa, 0, 0, \kappa$)

$$\langle k, \lambda | k', \lambda' \rangle = 2|k| \delta(\vec{k} - \vec{k}') \delta_{\lambda\lambda'}$$

b) $\lambda = \pm j$, j - integer



$$\Rightarrow S_{\pm j}(k, p) = \xi_{\pm}^{\pm j_1}(k) \dots \xi_{\pm}^{\pm j_j}(k_j) M_{\pm j_1, \dots, j_j}$$

$$k_{j_1} M_{j_1, \dots, j_j} = 0$$

$$\xi_{\pm}^{\pm j}(k) = R(k)^\wedge \bar{\xi}_{\pm}^{\nu}$$

$$\bar{\xi}_{\pm}^{\nu} = \frac{1}{\sqrt{2}} (0, 1, \pm i, 0)$$

c)

(2)

$$\begin{array}{c}
 p_i^\sigma \\
 \swarrow \\
 \text{---} \\
 \searrow \\
 p_i^{\sigma'}
 \end{array}
 \begin{array}{c}
 k \\
 \text{---} \\
 \text{---} \\
 \text{---}
 \end{array}
 \quad k \ll p$$

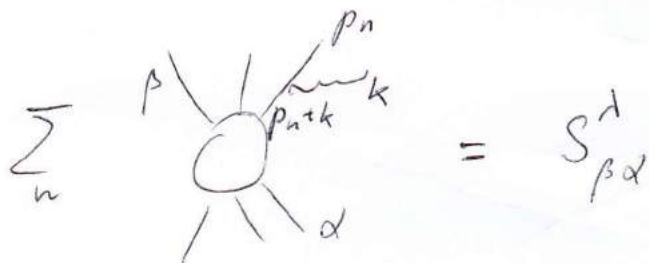
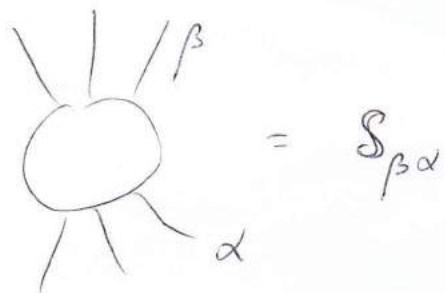
$$\propto \hat{\epsilon}_\pm^{\mu}(\vec{k}) - \hat{\epsilon}_\pm^{\nu}(\vec{k}) p_{\mu} - p_{\nu} \delta_{\sigma\sigma'}$$

N.B. Not necessarily diagonal in polarization σ, σ' at this stage, but see later

$$\text{photon} : \quad \frac{2ie_\pm \hat{\epsilon}_\pm^{\mu}(\vec{k}) p_{\mu}}{(2\bar{u})^{3/2}} \delta_{\sigma\sigma'} (2\bar{u})^4$$

$$\text{graviton} : \quad \frac{2i\kappa_\pm}{M_p} \frac{(p_{\mu} \hat{\epsilon}_\pm^{\mu}(\vec{k}))^2}{(2\bar{u})^{3/2}} \delta_{\sigma\sigma'} (2\bar{u})^4$$

2) Charge conservation



$$\text{out} : \quad \frac{1}{(p_n+k)^2 + m^2} \approx \frac{1}{2(p_n \cdot k)}$$

$$\text{in} : \quad \frac{1}{(p_n-k)^2 + m^2} \approx -\frac{1}{2(p_n \cdot k)}$$

$$S_{\beta\alpha}^{\pm 1} = \sum_n \frac{\eta_n}{2(p_n \cdot k)} \frac{(\tilde{\Sigma}_{\pm}^{\mu\nu} \cdot p_{\mu} p_{\nu}) \cdot 2i e_n (2\bar{n})^4}{(2\bar{n})^{3/2}} \left(-\frac{i}{(2\bar{n})^4}\right) S_{\beta\alpha}$$

$$\eta_n = \begin{cases} +1 & \text{- outgoing} \\ -1 & \text{- incoming} \end{cases}$$

$$S_{\beta\alpha}^{\pm 1} = \frac{S_{\beta\alpha}}{(2\bar{n})^{3/2}} \sum_n e_n \eta_n \frac{\tilde{\Sigma}_{\pm}^{\mu\nu}(\bar{k}) p_{\mu} p_{\nu}}{p_n \cdot k}$$

$$\tilde{\Sigma}^{\mu\nu} \rightarrow k^{\mu} \Rightarrow 0 = k_{\mu} M^{\mu\nu} \propto \sum_n e_n \eta_n$$

$$\Rightarrow \sum_{out} e_n \eta_n = \sum_{in} e_n \eta_n = 0$$

3) Equivalence principle

$$S_{\beta\alpha}^{\pm 2} = \frac{S_{\beta\alpha}}{(2\bar{n})^{3/2}} \sum_n \frac{f_n}{M_p} \eta_n \frac{(\tilde{\Sigma}_{\pm}^{\mu\nu}(\bar{k}) p_{\mu} p_{\nu})^2}{(p_n \cdot k)}$$

$$0 = \left(\sum_n \eta_n f_n p_{\mu} p_{\nu}\right) \tilde{\Sigma}_{\pm}^{\mu\nu} S_{\beta\alpha} + \text{no additional kinematic constraints}$$

\Rightarrow all f_n are equal $\Rightarrow \sum \eta_n p_{\mu} p_{\nu} = 0$ due to energy - momentum conservation

N.B. a) $f_+ \neq f_-$ in general (unless parity)

b) graviton can be also in the hard lines \Rightarrow ~~interact~~ self interaction of gravitons

4) Higher spins?

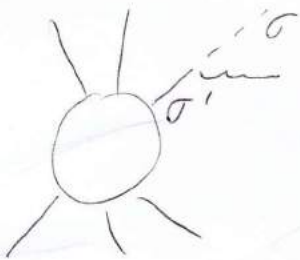
(4)

$$\sum_n g_n^{(j)} y_n p_{\mu_1} \dots p_{\mu_{j-1}} = 0$$

If no restrictions on kinematics $\Rightarrow g_n^{(j)} = 0$

Ex. 8 Does this argument exclude any interactions of massless higher spins?

Other couplings?



$$\sum_n S_{\beta\alpha}^{\text{spin-flipped}} \bar{e}_n y_n = 0$$

But spin-flipped amplitudes are different functions of the hard momenta

\Rightarrow cancellation is impossible

B. Maxwell equations

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5) Free massless fields
creation operators

$$a^\dagger(\vec{k}, \lambda) |0\rangle = |\vec{k}, \lambda\rangle$$

$$U(\Lambda) a^\dagger(k, \lambda) U^{-1}(\Lambda) = e^{i d\theta} a^\dagger(\Lambda k, \lambda)$$

$$U(\Lambda) a(k, \lambda) U^{-1}(\Lambda) = e^{-i d\theta} a(\Lambda k, \lambda)$$

definition of a local field $\psi_n(x)$

1. $\psi_n(x)$ - linear combination of a, a^\dagger

$$2. \quad i [P_\mu, \psi_n(x)] = \partial_\mu \psi_n(x)$$

3. $\psi_n(x)$ transforms in a finite-dimensional irrep of the Lorentz group

$$U(\Lambda) \psi_n(x) U^{-1}(\Lambda) = \sum_m D_{nm}(\Lambda) \psi_n(\Lambda x)$$

4. Locality: $[\psi_n(x), \psi_m(y)] = 0$ outside the light-cone

Needed to have Lorentz ~~covariant~~ invariant
time-ordering

First try: annihilation field

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$$\psi_n(x) = \frac{1}{(2\pi)^{3/2}} \int \frac{d^3k}{2|k|} e^{ikx} a(\vec{k}, t) \omega_n(\vec{k}, t)$$

$$e^{-i\Lambda\theta} (\mathcal{L}^{-1}(\Lambda k) \wedge \mathcal{L}(k)) \omega_n(\vec{k}, t) = \sum_m D_{nm}(\Lambda^{-1}) \omega_m(\vec{k}, t)$$

Take $k_f = \bar{k}_f$

$$\omega_m(\vec{k}, t) \equiv \bar{\omega}_m(t)$$

$$\Rightarrow e^{-i\Lambda\theta} (\mathcal{L}^{-1}(\Lambda k) \wedge \mathcal{L}(k)) \bar{\omega}_n = \sum_m D_{nm}(\Lambda^{-1}) \bar{\omega}_m$$

Let $\Lambda = L^{-1}$ belong to the little group

$$e^{+i\Lambda\theta(L)} \bar{\omega}_n = \sum_m D_{nm}(L) \bar{\omega}_m$$

$$L = 1 + i\alpha_3 J_3 + i\beta_1 (K_1 - J_2) + i\beta_2 (K_2 + J_2)$$

$$\theta = \alpha_3$$

$$\Rightarrow D_{nm}(J_3) \bar{\omega}_m = d \bar{\omega}_n$$

$$D_{nm}(K_1 - J_2) \bar{\omega}_m = D_{nm}(K_2 + J_2) \bar{\omega}_m = 0$$

$$J_i = N_i^+ + N_i^- , \quad K_i = -i(N_i^+ - N_i^-)$$

independent SU(2) generators

$$D_{nm} (N_1^+ - iN_2^+) \bar{w}_m = D_{nm} (N_1^- + iN_2^-) \bar{w}_m = 0$$

$$\Rightarrow D(N_3^+) \bar{w} = j_+ \bar{w} \quad ; \quad D(N_3^-) w = -j_- w$$

$$j_+ - j_- = 1$$

$\Rightarrow \bar{w}_n$ lies in the $(j_+ + \frac{1}{2}, j_-)$ representation

Basic rep. $(d, 0)$. The rest come from derivatives

Examples!

photons $\lambda = \pm 1 \Rightarrow \phi_\mu$ in $(1, 0)$ or $(0, 1)$
 $F_{\mu\nu}^+ \quad F_{\mu\nu}^-$

no $A_\mu \in (\frac{1}{2}, \frac{1}{2})$!

gravitons $\lambda = \pm 2 \Rightarrow (2, 0)$ or $(0, 2)$

Riemann tensor

no $g_{\mu\nu} \in (1, 1)$!

Explicitly for photons

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$$F_+^{\mu\nu} = i \int \frac{d^3k}{(2\pi)^{3/2} 2k} e^{ik\nu} (k^\mu \epsilon_+^\nu(\vec{k}) - k^\nu \epsilon_+^\mu(\vec{k})) a(k, +)$$

Ex. 9 Check that this field transforms as a (1,0) tensor

But this does not satisfy locality!

$$[F_+^{\mu\nu}(x), (F_+^{\rho\sigma}(y))^{\dagger}]$$

$$\propto \int \frac{d^3k}{2|k|} (k^\mu \epsilon_+^\nu - k^\nu \epsilon_+^\mu) (k^\rho \epsilon_+^{\sigma\dagger} - k^\sigma \epsilon_+^{\rho\dagger}) e^{ik(x-y)}$$

Take $x^0 = y^0$, $\mu = \rho = 0$, $\nu = \sigma = i$

$$\Rightarrow \propto \int \frac{d^3k}{2|k|} e^{i\vec{k}(\vec{x}-\vec{y})} |k|^2 \epsilon_+^i \epsilon_+^{*i} = \underbrace{\int \frac{d^3k}{2} |k| e^{i\vec{k}(\vec{x}-\vec{y})}}_{\neq 0}$$

We need opposite helicity

$a^\dagger(k, -)$ transforms in the same way as $a(k, +)$

$$\Rightarrow F_+^{\mu\nu} = i \int \frac{d^3k}{(2\pi)^{3/2} 2|k|} (k^\mu \epsilon_+^\nu - k^\nu \epsilon_+^\mu) (a(k, +) e^{ikx} + y a^\dagger(k, -) e^{-ikx})$$

Ex. 10 Show that $F_+^{\mu\nu}$ commutes with $(F_+^{\rho\sigma})^\dagger$ at space-like separations if $|y| = 1$

We choose $y = -1$

N.B. $(F_+^{\mu\nu})^\dagger = F_-^{\mu\nu}$

6) Interaction

(9)

Attempt to use $F_{\mu\nu}^{\pm}$

$$\text{But } \langle T [F_{\mu\nu}^{\pm}(k) F_{\rho\sigma}^{\pm}(-k)] \rangle \propto \frac{k_{\mu} k_{\nu} g_{\rho\sigma} + \dots}{k^2}$$

\Rightarrow no pole at $k^2 \rightarrow 0 \Rightarrow$ no long-range force

$$\text{Notice } F_{\pm}^{\mu\nu} = \partial_{\mu} A_{\pm\nu} - \partial_{\nu} A_{\pm\mu}$$

$$A_{\pm}^{\mu} = \int \frac{d^3k}{(2\pi)^{3/2} 2k} \epsilon_{\pm}^{\mu}(k) (a(k,+)e^{ikx} + a^{\dagger}(k,-)e^{-ikx})$$

$$A_{\pm}^0 = 0 \Rightarrow A_{\pm}^{\mu} \text{ are } \underline{\text{not}} \text{ vectors}$$

Indeed, we have seen that it is impossible to have $(\frac{1}{2}, \frac{1}{2})$ field for photons

Ex. 11 Show that

$$U(\Lambda) A_{\pm}^{\mu}(x) U(\Lambda)^{-1} = \Lambda_{\nu}^{\mu} A_{\pm}^{\nu}(\Lambda x) + \partial_{\mu} P_{\pm}(x, \Lambda)$$

Real combinations

$$A_{\mu} = A_{+\mu} + A_{-\mu} \quad - \text{vector}$$

$$B_{\mu} = -i(A_{+\mu} - A_{-\mu}) \quad - \text{pseudo-vector} \quad \left. \begin{array}{l} \text{under} \\ \text{P and T} \end{array} \right\}$$

$$\Delta_{\mu\nu}(k) = \langle T(A_\mu(k) A_\nu(-k)) \rangle =$$

$$= -i \sum_{S=\pm} \frac{(\epsilon_S^\mu \epsilon_S^\nu)(\vec{k})}{k^2 - i\epsilon} = -\frac{i}{k^2 - i\epsilon} \left(\eta_{\mu\nu} + \frac{k_1^\mu k_2^\nu + k_1^\nu k_2^\mu}{2|\vec{k}|^2} \right)$$

$\Pi^{\mu\nu}$ from lecture 1

$$k_1^\nu = (|\vec{k}|, \vec{k})$$

$$k_2^\nu = (|\vec{k}|, -\vec{k})$$

$$\Pi^{\mu\nu} = \eta^{\mu\nu} + \underbrace{(n^\mu k^\nu + n^\nu k^\mu)}_{\text{grad}} \frac{k^0}{|\vec{k}|^2} - \frac{k^\mu k^\nu}{|\vec{k}|^2} + \underbrace{\frac{k^2 n^\mu n^\nu}{|\vec{k}|^2}}_{\text{time-like}}$$

$$n^\mu = (1, 0, 0, 0)$$

grad

time-like

$$\Delta_{\mu\nu}(k) = \underbrace{\frac{\eta_{\mu\nu}}{k^2}}_{\text{Lorentz covariant}} + \underbrace{\frac{1}{k^2} (k^\mu \tilde{n}^\nu + \tilde{n}^\nu k^\mu)}_{\Delta_{\text{grad}}^{\mu\nu}} + \underbrace{\frac{n^\mu n^\nu}{|\vec{k}|^2}}_{\text{no physical pole} \Leftrightarrow \text{instantaneous interaction } \Delta_{\text{inst}}^{\mu\nu}}$$

Lorentz covariant

$\Delta_{\text{grad}}^{\mu\nu}$

no physical pole \Leftrightarrow instantaneous interaction $\Delta_{\text{inst}}^{\mu\nu}$

$$H_{\text{int}} = \int d^3x (-A_i J^i + \dots)$$

~~Angular~~

Non-covariant contributions must cancel

$\Delta_{\text{grad}}^{\mu\nu}$

vanishes if J^μ is conserved, $\partial_\mu J^\mu = 0$

$\Delta_{\text{inst}}^{\mu\nu}$

cancelled by an explicit non-local term:

$$H_{\text{int}} = \int d^3x (-A_i J^i) + \frac{1}{2} \int d^3x d^3y J^0(\vec{x}, t) D(\vec{x} - \vec{y}) J^0(\vec{y}, t)$$

$$D(\vec{x}) = \frac{1}{(2\pi)^3} \int \frac{d^3k}{|k|^2} e^{ikx} = \frac{1}{4\pi |\vec{x} - \vec{y}|} \quad (1)$$

Convenient to introduce $A_0(x, t)$ with the propagator

$$\langle A_0(x, t), A_0(x', t') \rangle = \frac{\delta(t-t')}{4\pi |x-x'|}$$

\Rightarrow Recover QED in the Coulomb gauge

Ex. 12 Assume that the theory also contains other fields charged under B_μ . What is the physical interpretation of such interaction?

Compute the propagator

$$\langle T(A_\mu(x) B_\nu(-k)) \rangle$$

Is it possible to decompose it into Lorentz-covariant, gradient and instantaneous parts?

Analyse physical implications of the result.